



XII ESCOLA DO CBPF

22 de julho a 02 de agosto de 2019

Inteligência Artificial Utilizando *Deep Learning* e Aplicações em Física



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(CBPF)**



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(CBPF)**

Uma introdução conceitual de aprendizado de máquina e uma abordagem prática em aplicações de redes neurais profundas com foco em aplicações científicas e tecnológicas.





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Ementa

AULA 1: Apresentações iniciais, Redes Neurais Artificiais, Aprendizado de máquina, Algoritmos/Termos; Redes Perceptron de Várias Camadas.

AULA 2: Rede Neural Convolucional: Treinamento, Maxpooling, Dropout

AULA 3: Overfitting, funções de ativação. Redes VGG e AlexNet, Resnet e Inception

AULA 4: Autoencoders e aprendizagem não supervisionada.
- Redes Geradoras Adversariais (GAN).

AULA 5: Introdução a Redes Neurais Bayesianas, Redes Neurais Convolucionais Baseadas em Região (R-CNN). Introdução a Reinforcement Learning

Exemplos / Demonstrações

Pré-requisitos:

Recomenda-se o domínio da linguagem de programação Python
Porém é possível o uso de C ou MATLAB.



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You will need:

Python (anaconda3)

Keras

Tensorflow

matplotlib

numpy

For the examples you will also need:

astropy (Aulas Clécio)

Google Colab (on-line)

We will have several examples in Google Colabs or in Python notebooks. The examples and datasets required can be downloaded in



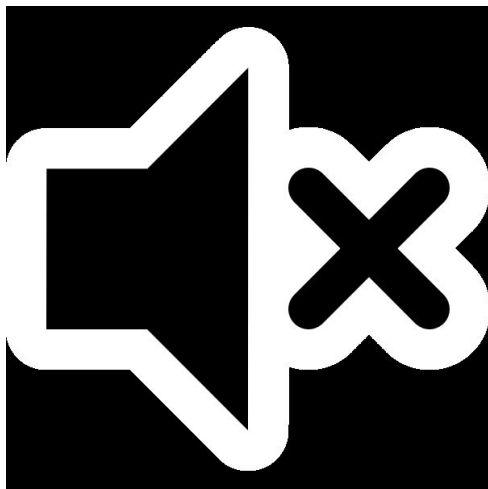
[Clearnightsthebest.com](https://clearnightsthebest.com)

debom@cbpf.br



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22 de julho a 02 de agosto de 2019



Para um melhor aproveitamento do conteúdo ministrado, recomendamos à audiência que mantenha os aparelhos eletrônicos (celulares, laptops) Desligados ou mudos durante as aulas.



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login: XII-Escola-CBPF
senha: escolacbpf2019

Laboratório de Redes e Sistemas (LARS)



6.RTX2080TI

36x RTX 2080Ti
6x Quadro K6000
6x 1080Ti
13x 1050Ti

- 32 CPU cores, 128 GB RAM (expansível até 1TB),
- 26112 CUDA cores, 84 Teraflops, 66GB VRAM,
- 4xWaterCoolers - completo silêncio,
- 1632 Tmus unid. de texturas e 528 ROps unid. para renderização
- 4TB HD/backup Enterprise e 500 GB SSD/OS, 2 fontes 1600W



GPGPU – General Purpose GPU (processador da placa de vídeo)



As GPU foram originalmente usadas para renderização de imagens em jogos 3D.

Essa capacidade de processamento está sendo aproveitada de forma mais ampla para acelerar o processamento computacional em áreas como pesquisa científica, processamento de sinais/imagens, algebra linear, estatística, reconstrução 3D, modelagem financeira, exploração de petróleo e gás....

Laboratório de Redes e Sistemas (LARS) Programação Paralela



SciMining x Santos Dumont (2) PERFORMANCE

FP32
(float)

HPC	Modelo	TFLOPS	Nós	GPU/ Nós	Total TFLOPS		Percent ual
SDU	K40	5,046	198	2	1998,22		100%
SciMining	RTX2080Ti	13,45	1	6	80,70		4%
SciMining 7	RTX2080Ti	13,45	7	6	564,90		28%

[https://www.tec
hpowerup.com/g
pu-specs/geforce
-rtx-2080-ti.c330](https://www.tec
hpowerup.com/g
pu-specs/geforce
-rtx-2080-ti.c330)

Laboratório de Redes e Sistemas (LARS) Programação Paralela



Fermilab (FP32 (float) performance)			
K20	2	2,87	5,74
K40	2	4,29	8,58
P100	10	9,526	95,26
V100	4	14,13	56,52
Total			166,1

<https://wilsonweb.fnal.gov/phigpu.shtm>

↓

Laboratório de Redes e Sistemas (LARS) Programação Paralela

SciMining x Wilson(3) - RENDER CONFIG



		TENSOR CORES					
HPC	Modelo	TENSOR CORES	Nós	GPU/ Nós	Total TENSOR CORES		Percent ual
Wilson	Varios	2560			2560		100%
SciMIning	RTX2080Ti	544	1	6	3264		128%
SciMIning 7	RTX2080Ti	544	7	6	22848		700%

<https://www.techpowerup.com/gpu-specs/geforce-rtx-2080-ti.c3305>



GPU – Unidade de Processamento Gráfico (processador da placa de vídeo)

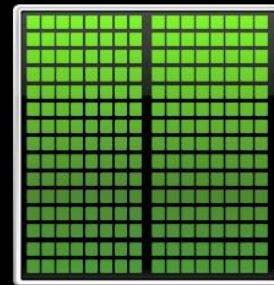
A GPU é responsável pelos cálculos de processamento de vetores, texturas, projeções das imagens em um sistema de vídeo do computador.

Libera a CPU para fazer outras tarefas (de outra complexidade).

The Difference between a CPU and GPU



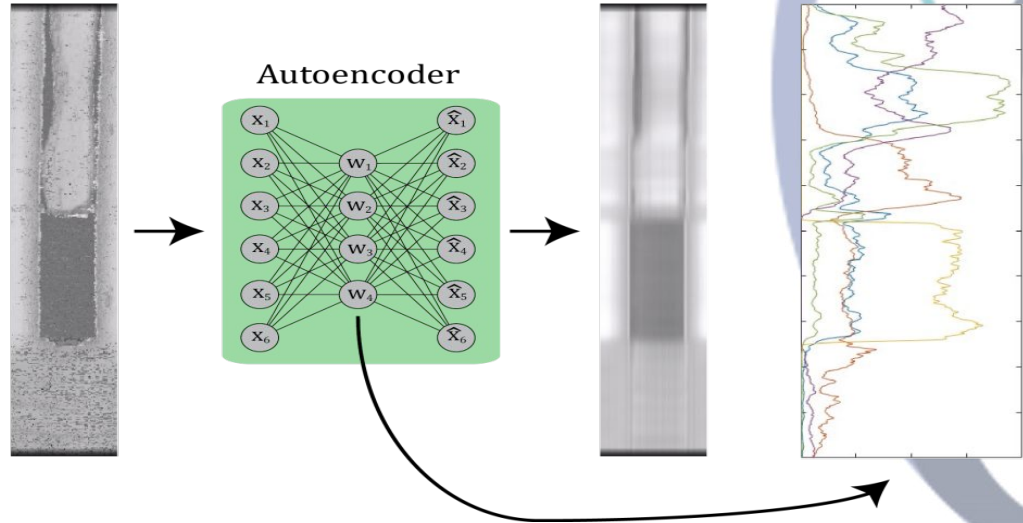
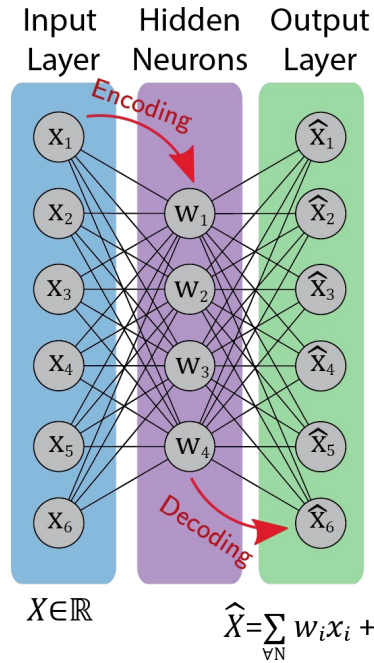
CPU



GPU

What is an AutoEncoder?

- ANN aim to reproduce input data.
- **Encode/Decode** strategy.
- **IDENTITY FUNCTION** → Hidden Features



Auto Encoders Example

```
from keras.layers import Input, Dense
from keras.models import Model

# this is the size of our encoded representations
encoding_dim = 32
# 32 floats -> compression of factor 24.5, assuming the input is 784
floats

# this is our input placeholder
input_img = Input(shape=(784,))
# "encoded" is the encoded representation of the input
encoded = Dense(encoding_dim, activation='relu')(input_img)
# "decoded" is the lossy reconstruction of the input
decoded = Dense(784, activation='sigmoid')(encoded)

# this model maps an input to its reconstruction
autoencoder = Model(input_img, decoded)
```

Auto Encoders – A simple example

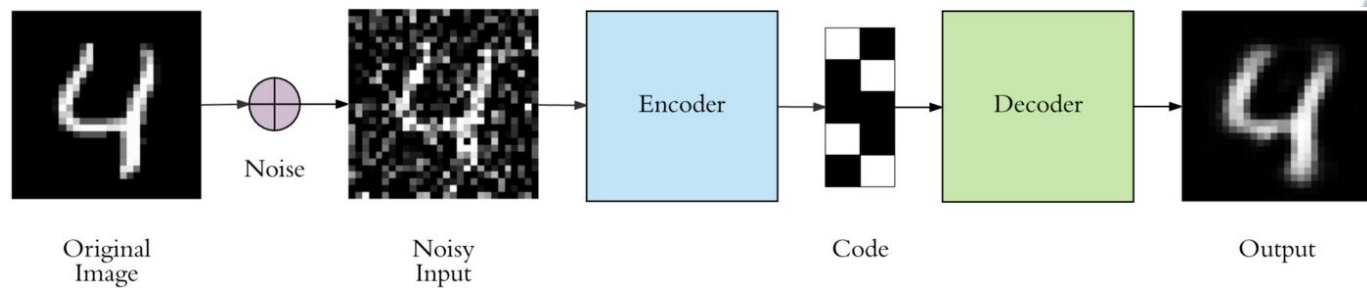
```
# this model maps an input to its encoded representation
encoder = Model(input_img, encoded)

# create a placeholder for an encoded (32-dimensional) input
encoded_input = Input(shape=(encoding_dim,))
# retrieve the last layer of the autoencoder model
decoder_layer = autoencoder.layers[-1]
# create the decoder model
decoder = Model(encoded_input, decoder_layer(encoded_input))
```

Warning: The autoencoders do overfit a lot if they are too complex.

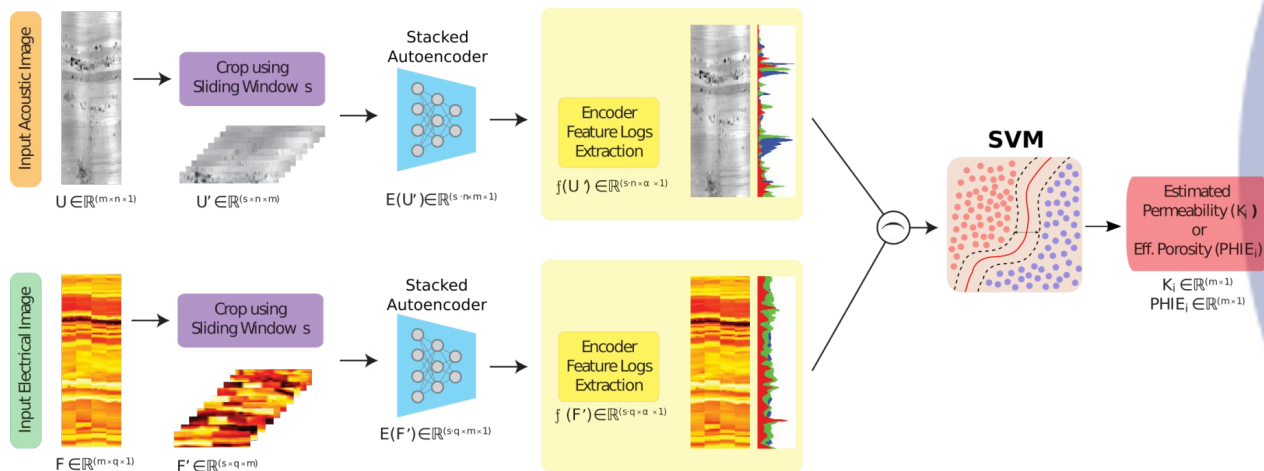
Small number of parameters forces the autoencoder to learn an intelligent representation of the data.

Denoising

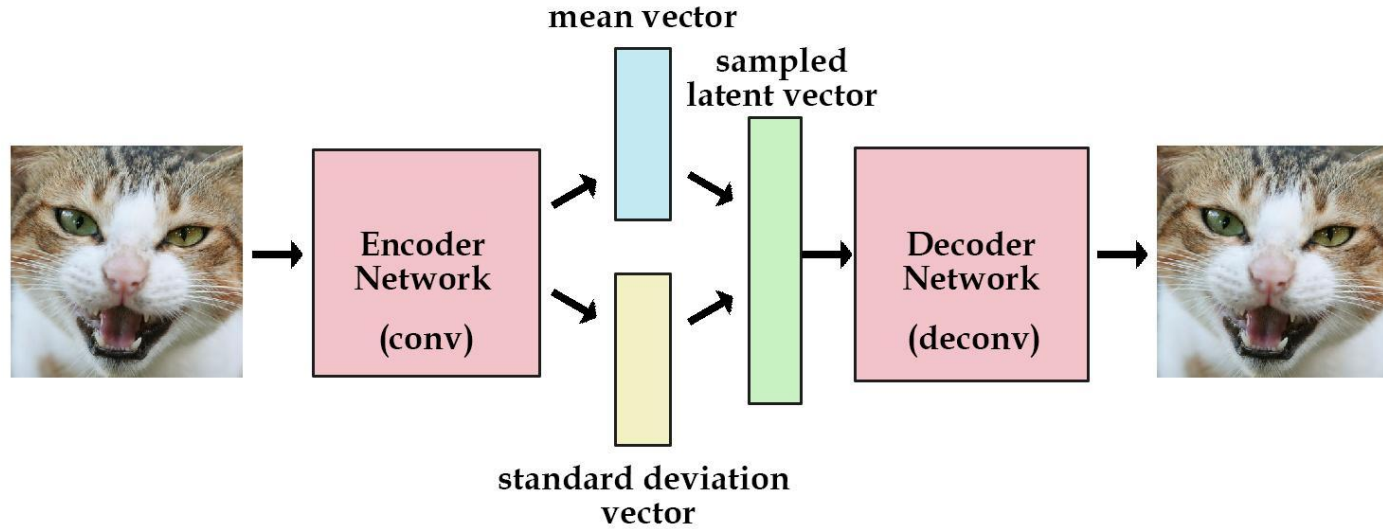


Regression with AE? But Why?

- Small datasets
- “Empty” datasets



Variational AE



Variational AE – Generative Model

An autoencoder that learns a latent variable model for its input data. So instead of letting your neural network learn an arbitrary function, you are learning the parameters of a probability distribution modeling your data. If you sample points from this distribution, you can generate new input data samples: a VAE is a "generative model".

$$q(Z|X)$$

Encode

$$p(X|Z)$$

Decode

```
x = Input(batch_shape=(batch_size, original_dim))  
h = Dense(intermediate_dim, activation='relu')(x)  
z_mean = Dense(latent_dim)(h)  
z_log_sigma = Dense(latent_dim)(h)
```

Variational AE – Generative Model

Latent normal distribution that is assumed to generate the data, via $\mathbf{z} = \mathbf{z_mean} + \exp(\mathbf{z_log_sigma}) * \epsilon$, where ϵ is a random normal tensor.

A decoder network maps these latent space points back to the original input data.

```
def sampling(args):  
    z_mean, z_log_sigma = args  
    epsilon = K.random_normal(shape=(batch_size,  
    latent_dim),  
                               mean=0.,  
                               std=epsilon_std)  
    return z_mean + K.exp(z_log_sigma) * epsilon  
  
z = Lambda(sampling,  
           output_shape=(latent_dim,))([z_mean, z_log_sigma])
```

$$q_{\phi}(\mathbf{z}_n | \mathbf{x}_n) = \mathcal{N}(\mathbf{z}_n | \mu_{\phi}(\mathbf{x}_n), \text{diag}(\sigma_{\phi}^2(\mathbf{x}_n))),$$

$$\mathbf{z} = g_{\phi}(\mathbf{x}, \epsilon) = \mu_{\phi}(\mathbf{x}) + \sigma_{\phi}(\mathbf{x}) \odot \epsilon, \quad \epsilon \sim \mathcal{N}(\mathbf{0}, \mathbf{I}),$$

Variational AE – Generative Model

Latent normal distribution that is assumed to generate the data, via $z = z_mean + \exp(z_log_sigma) * \epsilon$, where ϵ is a random normal tensor.

A decoder network maps these latent space points back to the original input data.

```
decoder_h = Dense(intermediate_dim, activation='relu')
decoder_mean = Dense(original_dim, activation='sigmoid')
h_decoded = decoder_h(z)
x_decoded_mean = decoder_mean(h_decoded)
```

Variational AE – Generative Model

Latent normal distribution that is assumed to generate the data, via $z = z_mean + \exp(z_log_sigma) * \epsilon$, where epsilon is a random normal tensor.

A decoder network maps these latent space points back to the original input data.

```
# end-to-end autoencoder
vae = Model(x, x_decoded_mean)

# encoder, from inputs to latent space
encoder = Model(x, z_mean)

# generator, from latent space to reconstructed inputs
decoder_input = Input(shape=(latent_dim,))
_h_decoded = decoder_h(decoder_input)
_x_decoded_mean = decoder_mean(_h_decoded)
generator = Model(decoder_input, _x_decoded_mean)
```

Variational AE – Generative Model

We add an extra term to the Loss

$$D_{\text{KL}}(Q\|P) = \sum_i Q(i) \ln\left(\frac{Q(i)}{P(i)}\right)$$

Consider an observed variable that is connected to the latent variable z .
As we are dealing with gaussians we may write

$$\text{KL}[q_\phi(\mathbf{z}|\mathbf{x})\|p(\mathbf{z})] = -\frac{1}{2} \sum_{k=1}^K \{1 + \log \sigma_k^2 - \mu_k^2 - \sigma_k^2\}$$

D. P. Kingma and M. Welling, "Auto-Encoding Variational Bayes," in Proceedings of the 2nd International Conference on Learning Representations (ICLR), 2014.

<http://louisiao.me/posts/implementing-variational-autoencoders-in-keras-beyond-the-quickstart-tutorial/#kingma2014>

Variational AE – Generative Model

```
def vae_loss(x, x_decoded_mean):  
    xent_loss = objectives.binary_crossentropy(x,  
    x_decoded_mean)  
    kl_loss = - 0.5 * K.mean(1 + z_log_sigma -  
    K.square(z_mean) - K.exp(z_log_sigma), axis=-1)  
    return xent_loss + kl_loss
```

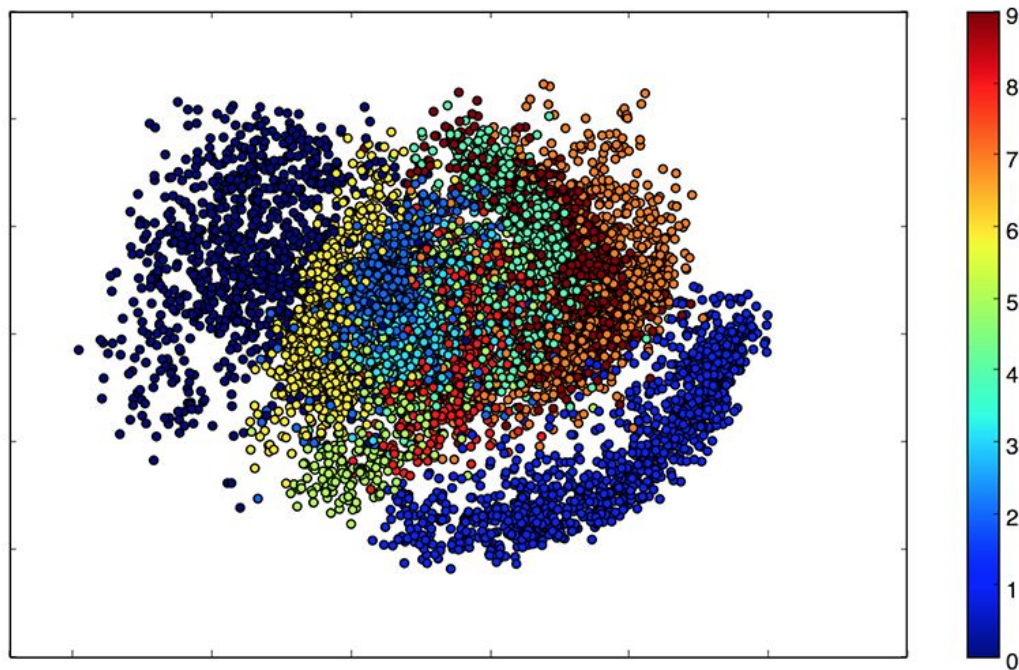
Variational AE – Generative Model

```
(x_train, y_train), (x_test, y_test) = mnist.load_data()

x_train = x_train.astype('float32') / 255.
x_test = x_test.astype('float32') / 255.
x_train = x_train.reshape((len(x_train),
                             np.prod(x_train.shape[1:])))
x_test = x_test.reshape((len(x_test), np.prod(x_test.shape[1:])))

vae.fit(x_train, x_train,
        shuffle=True,
        epochs=epochs,
        batch_size=batch_size,
        validation_data=(x_test, x_test))
```

```
x_test_encoded = encoder.predict(x_test, batch_size=batch_size)
plt.figure(figsize=(6, 6))
plt.scatter(x_test_encoded[:, 0], x_test_encoded[:, 1], c=y_test)
plt.colorbar()
plt.show()
```



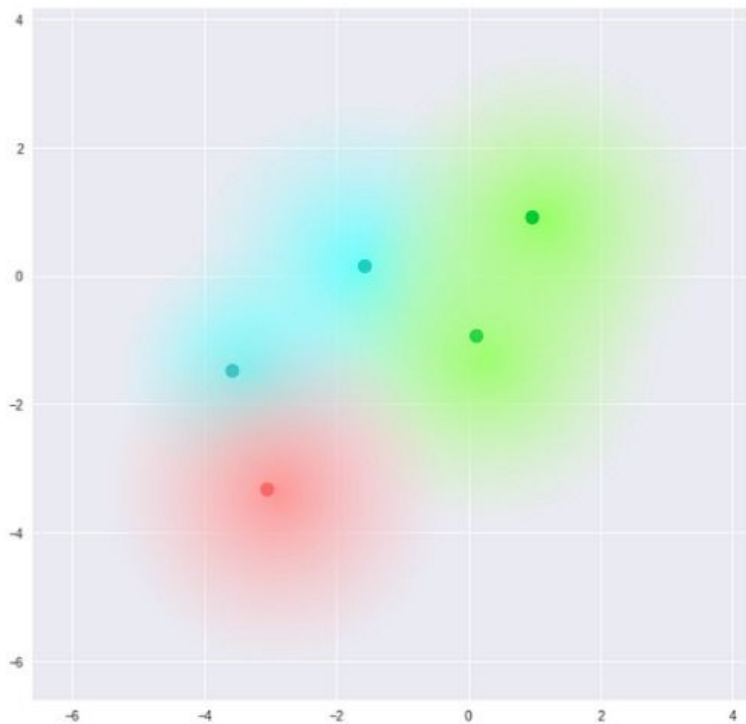
Variational AE – Generative Model

```
n = 15 # figure with 15x15 digits
digit_size = 28
figure = np.zeros((digit_size * n, digit_size * n))
# we will sample n points within [-15, 15] standard deviations
grid_x = np.linspace(-15, 15, n)
grid_y = np.linspace(-15, 15, n)

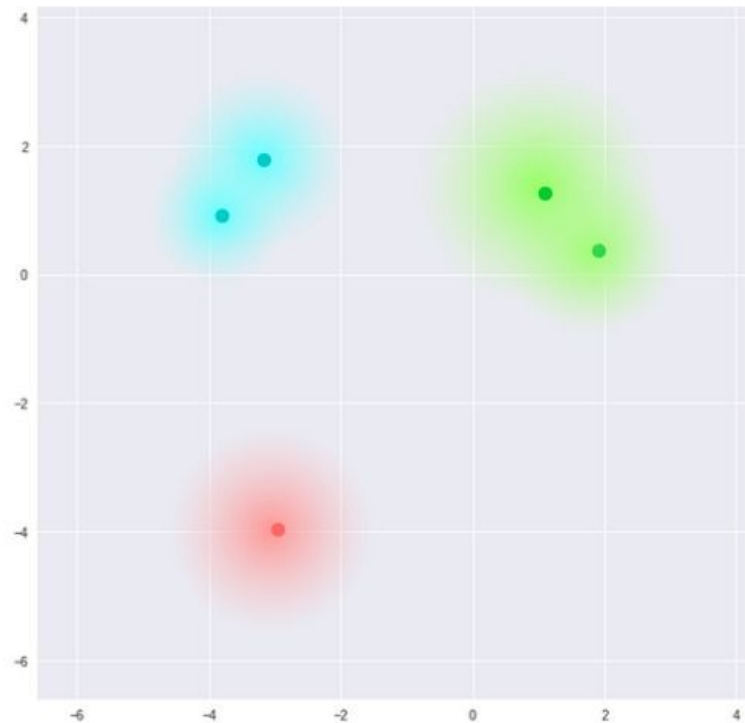
for i, yi in enumerate(grid_x):
    for j, xi in enumerate(grid_y):
        z_sample = np.array([[xi, yi]]) * epsilon_std
        x_decoded = generator.predict(z_sample)
        digit = x_decoded[0].reshape(digit_size, digit_size)
        figure[i * digit_size: (i + 1) * digit_size,
              j * digit_size: (j + 1) * digit_size] = digit

plt.show()
```

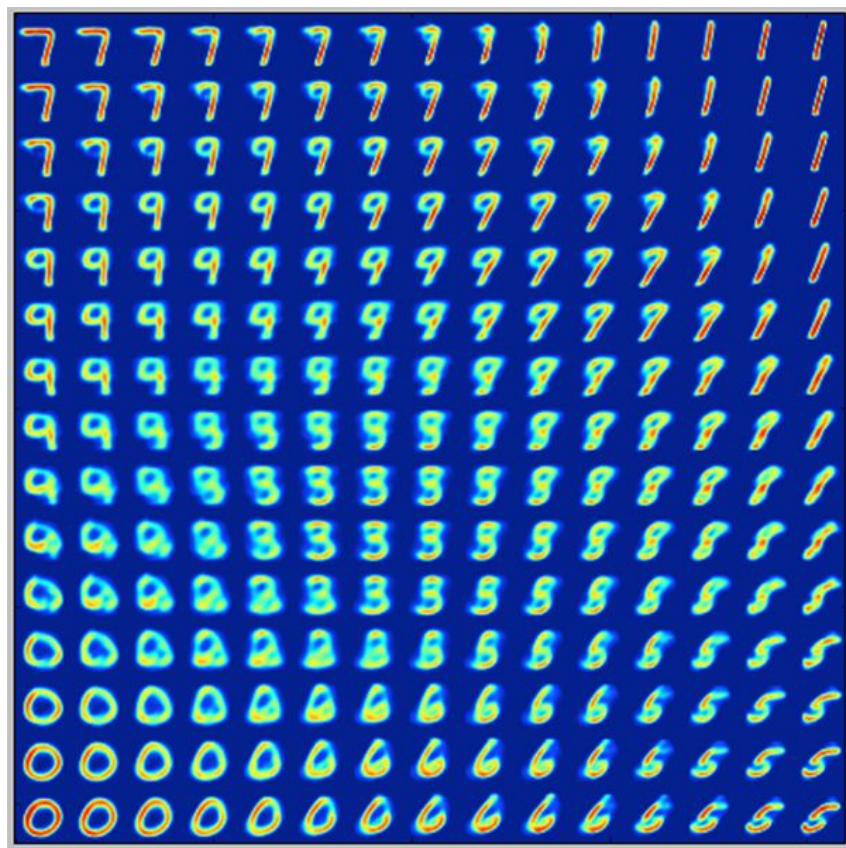
Variational AE – Generative Model



What we require



What we may inadvertently end up with



PCA vs AE

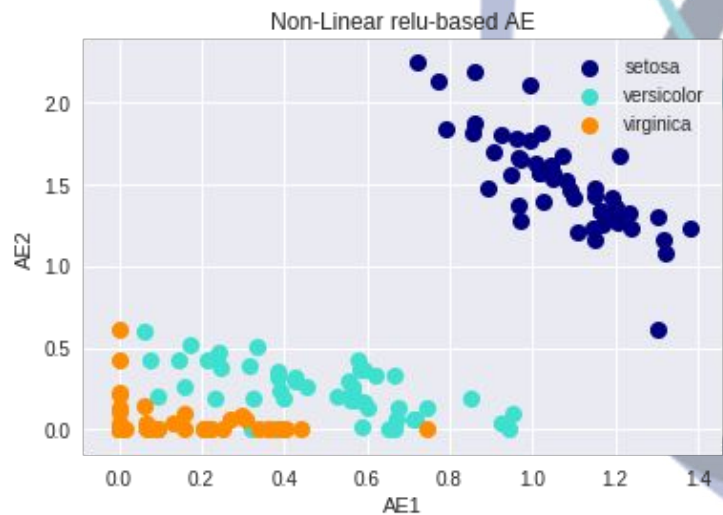
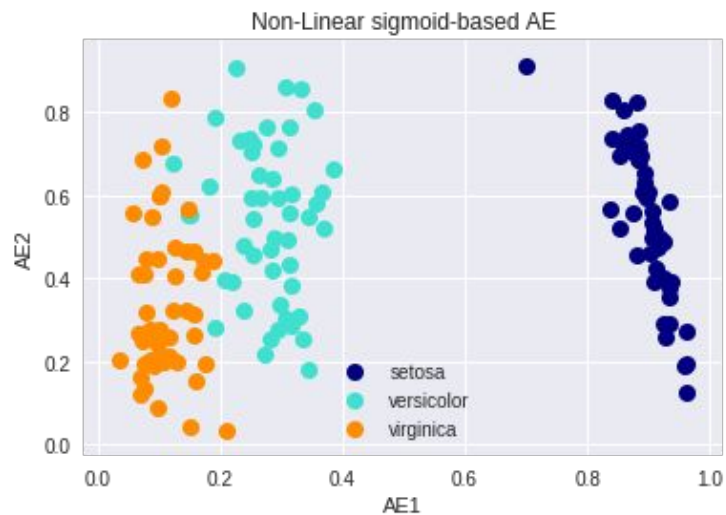
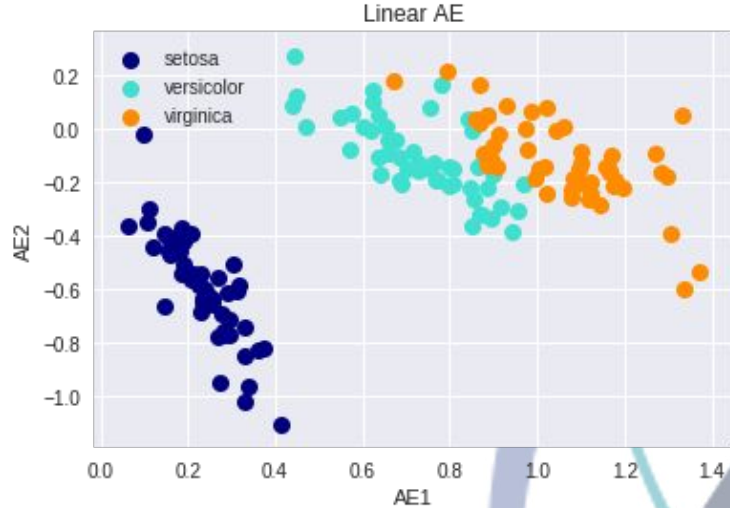
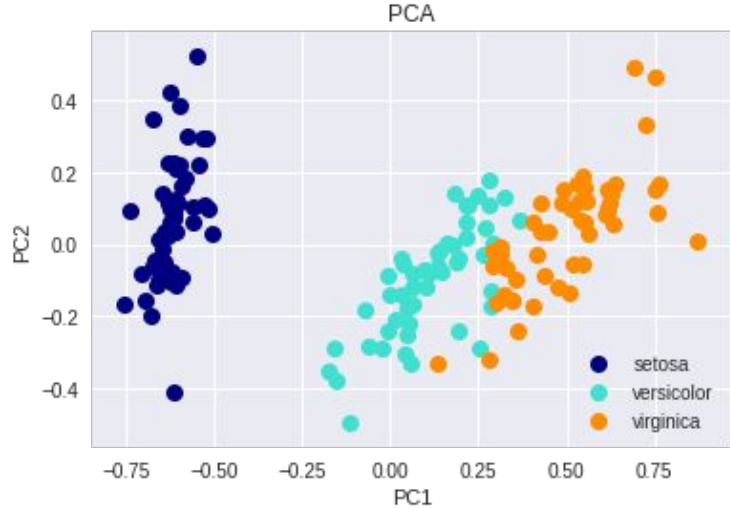
Iris Dataset

Attribute Information:

1. sepal length in cm
2. sepal width in cm
3. petal length in cm
4. petal width in cm
5. class:
 - Iris Setosa
 - Iris Versicolour
 - Iris Virginica

```
iris = datasets.load_iris()
X = iris.data
y = iris.target
target_names = iris.target_names

pca = decomposition.PCA()
pca_transformed = pca.fit_transform(X_scaled)
plot3clusters(pca_transformed[:, :2], 'PCA', 'PC')
```

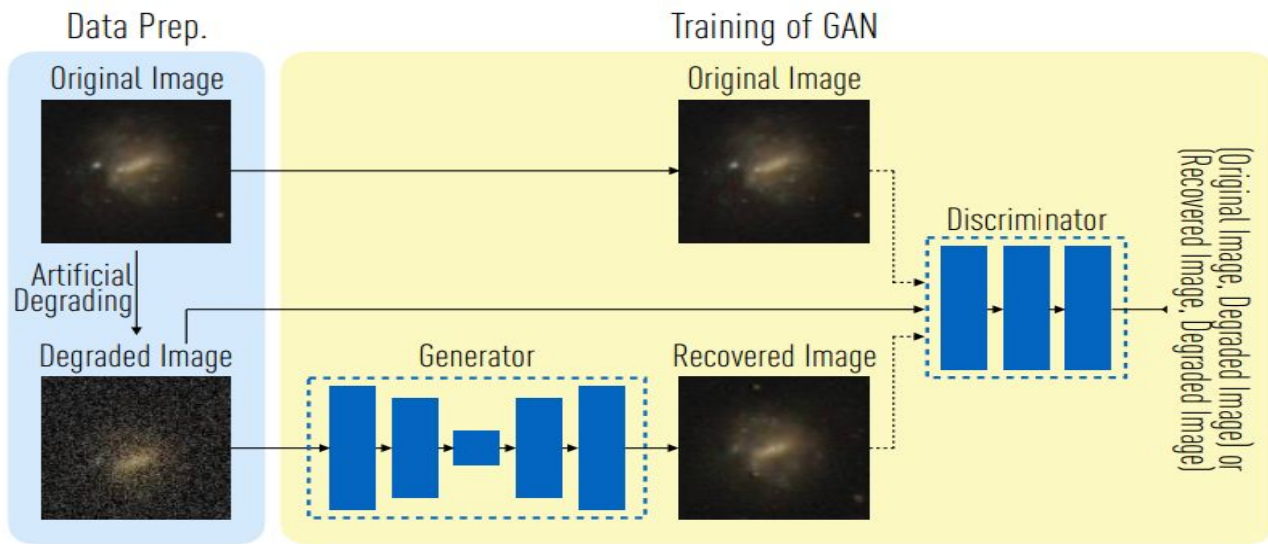


Generative Adversarial Networks

First proposed by Goodfellow et al. 2014 arXiv:1406.2661

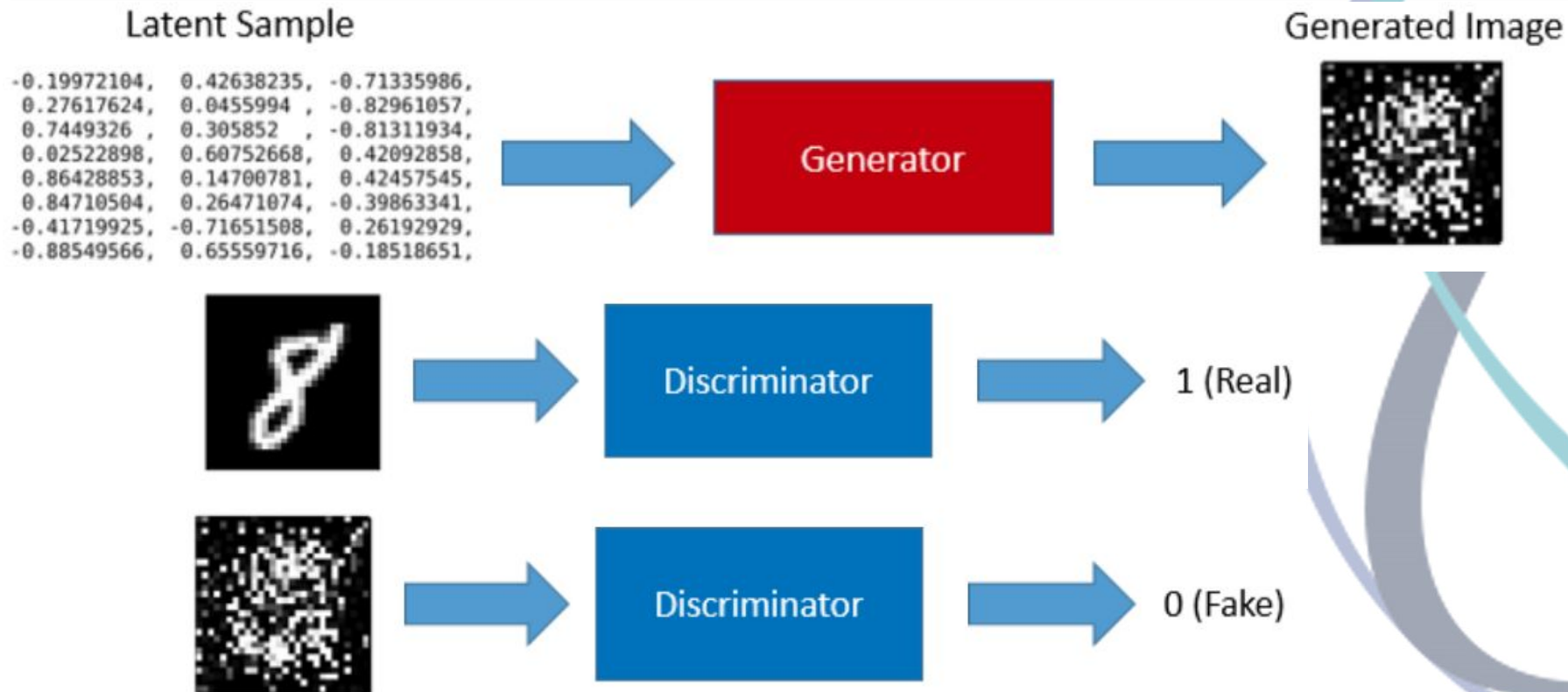
The Main idea is to make two networks to be adversaries.

(arXiv:1702.00403v1).



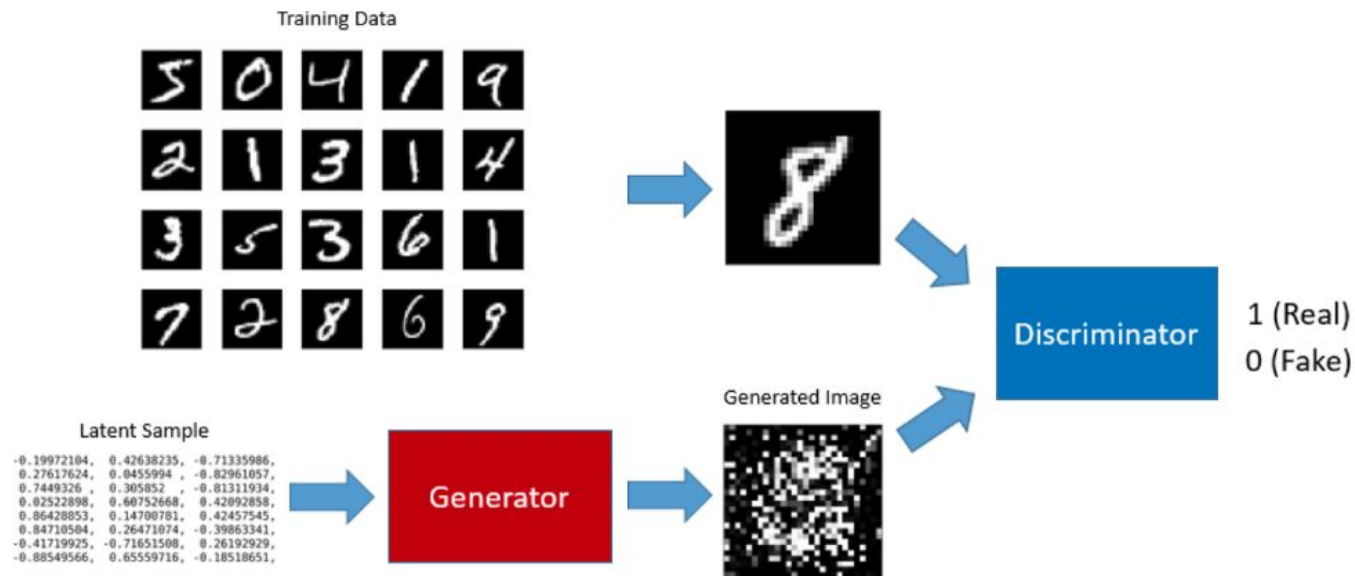
Generative Adversarial Networks

One need to set two networks: Generator and a Discriminator



Generative Adversarial Networks

The key here is the training There is a loop:
first we train the discriminator. Let the generator create some images out of noise and train **the discriminator only** to define true or false.



Generative Adversarial Neural Networks

The key here is the training There is a loop:

first we train the discriminator. Let the generator create some images out of noise and train **the discriminator only** to define true or false. Consider a Keras model D to be the discriminator, G the Generator and AD the Adversarial Model, i.e., the full model [G,D]

```
generated_images = G.predict(x=random_data, batch_size=batch_size)

d_loss_real = D.train_on_batch(image_real_batch, output_true)
d_loss_fake = D.train_on_batch(generated_images, output_false)
d_loss = 0.5 * np.add(d_loss_fake, d_loss_real)
```

Generative Adversarial Neural Networks

The key here is the training

Then we set the Discriminator to be **non-trainable**. After that we optimize the generator. We create some random noise data flows through the Network (Generator and Discriminator). The Generator weights will be updated by backpropagation from the non-trainable Discriminator to the Generator.

```
D.trainable = False
y = np.ones([batch_size, 1])
noise = np.random.uniform(-1.0, 1.0, size=[batch_size, 100])
AD_loss = AD.train_on_batch(random_data, y)
D.trainable = True
```



Generative Adversarial Neural Networks are hard to train...

Some general Advice

Gain is particularly sensible to vanishing gradients/ sparse gradients

LeakyReLU is recommended instead of ReLU.

If Discriminator loss converges rapidly and thus prevents the Generator from learning make its learning rate bigger than the Adversarial model learning rate you can also use a different training noise sample for the Generator.

If generated images look like noise. Use dropout on both Discriminator and Generator. Low dropout values generate more realistic images.



Some Scary smiles from DCGAN



Cycle GAN: Make your own Monet

[arXiv:1703.10593](https://arxiv.org/abs/1703.10593)



Monet $\leftrightarrow$ Photos



Monet $\rightarrow$ photo

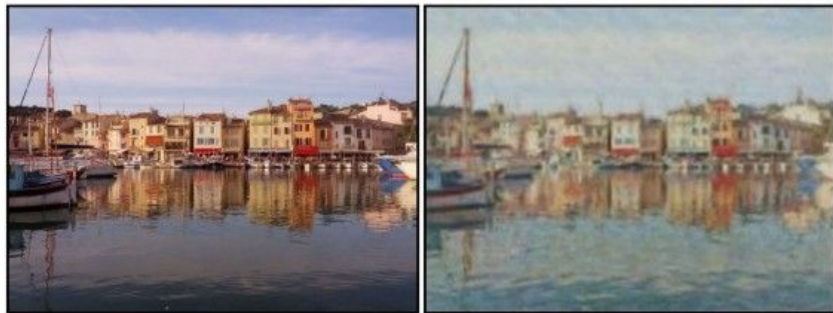
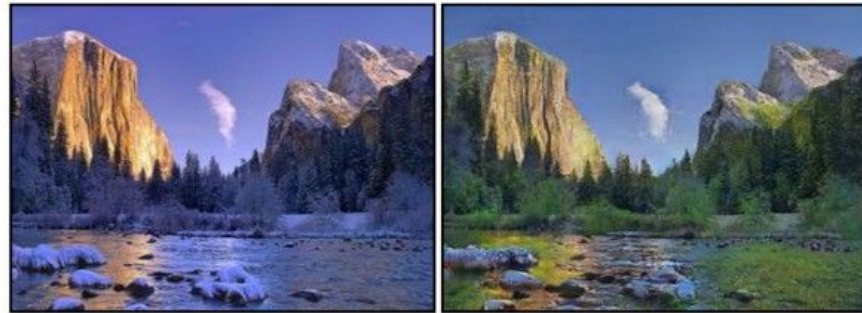


photo $\rightarrow$ Monet

Summer $\leftrightarrow$ Winter



summer $\rightarrow$ winter



winter $\rightarrow$ summer

Cycle GAN: Make your own Monet

[arXiv:1703.10593](https://arxiv.org/abs/1703.10593)



What people are doing with this?

DeepCMB: Lensing Reconstruction of the Cosmic Microwave Background with Deep Neural Networks

J. Caldeira^{a,1,*}, W. L. K. Wu^b, B. Nord^{b,c,e}, C. Avestruz^{a,b}, S. Trivedi^d, K. T. Story^f



[arXiv:1810.01483](https://arxiv.org/abs/1810.01483)



Deblending galaxy superpositions with branched generative adversarial networks

David M. Reiman^{1*†}, Brett E. Göhre^{1‡§}

¹Department of Physics, University of California Santa Cruz, 1156 High Street, Santa Cruz, California 95064, USA

[arXiv:1810.10098](https://arxiv.org/abs/1810.10098)

Enabling Dark Energy Science with Deep Generative Models of Galaxy Images

Siamak Ravanbakhsh¹, François Lanusse², Rachel Mandelbaum², Jeff Schneider¹, and Barnabás Póczos¹



[arXiv:1609.05796](https://arxiv.org/abs/1609.05796)



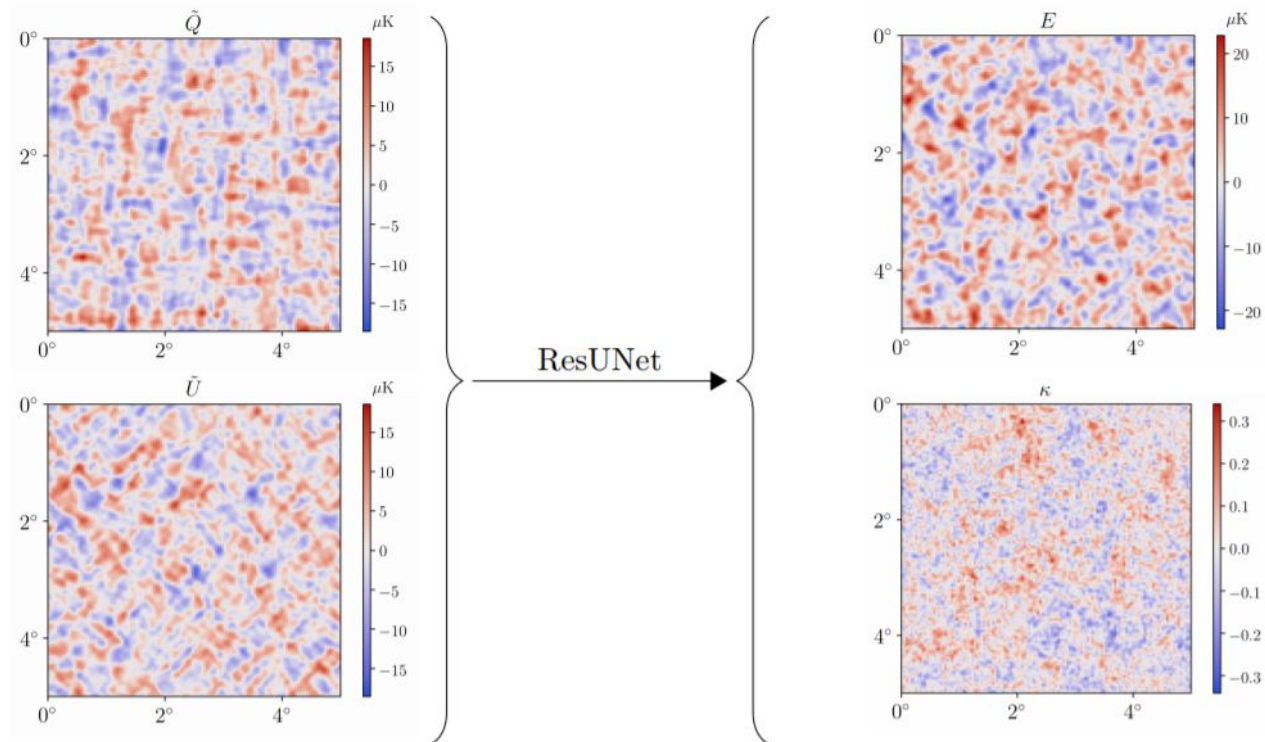
Fast Cosmic Web Simulations with Generative Adversarial Networks

A.C. Rodríguez,^a T. Kacprzak,^b A. Lucchi,^c A. Amara,^b R. Sgier,^b J. Fluri,^b T. Hofmann,^c and A. Réfrégier^b

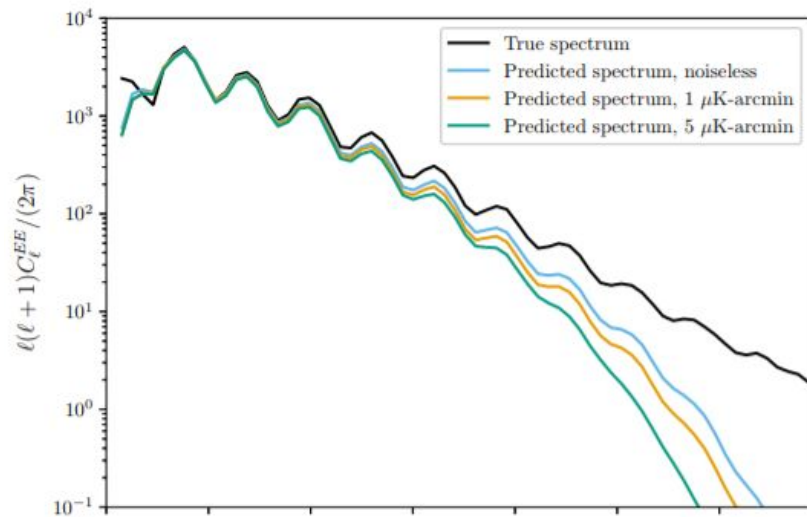
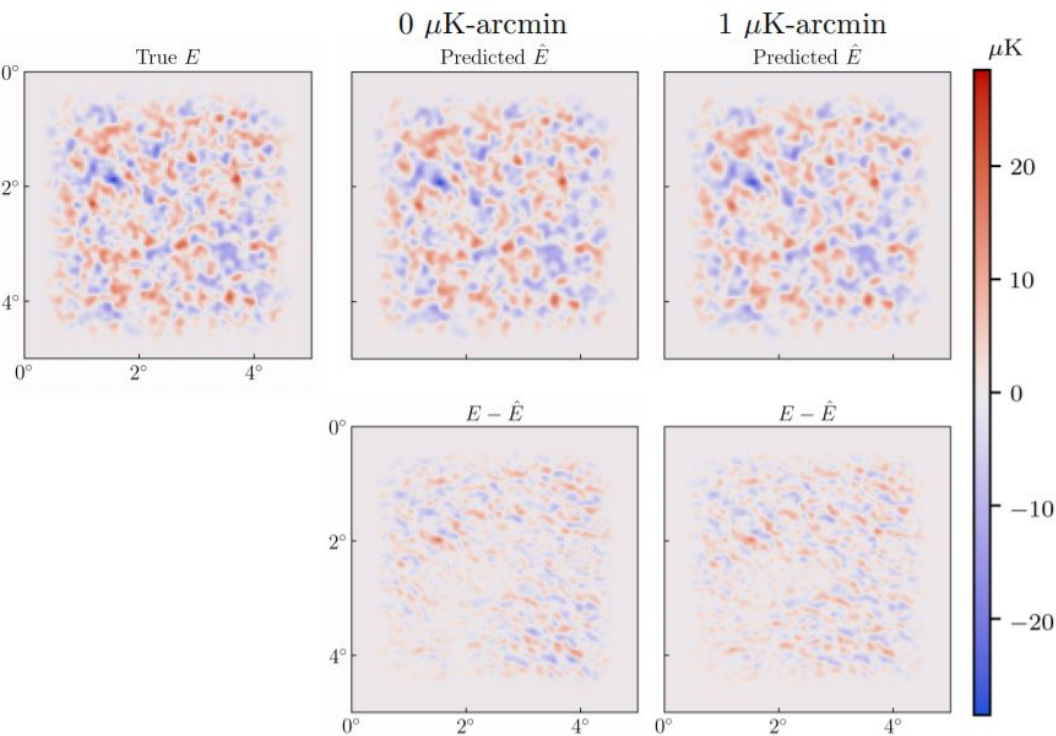
[arXiv:1801.09070](https://arxiv.org/abs/1801.09070)

DeepCMB: Lensing Reconstruction of the Cosmic Microwave Background with Deep Neural Networks

CMB polarization maps are usually represented (Q, U) basis corresponds to Stokes parameters and is convenient for mapping onto from the CMB instruments' polarization detector coordinates.

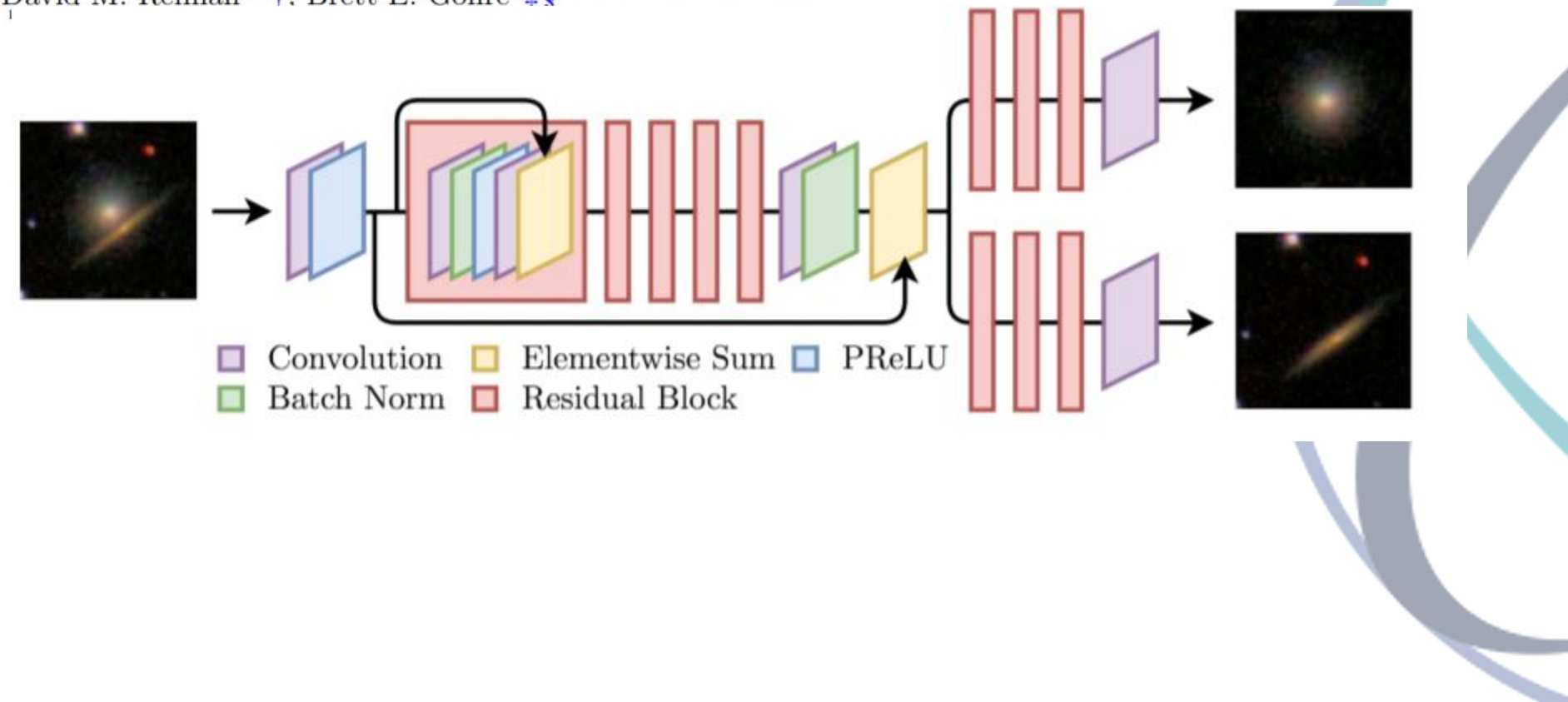


DeepCMB: Lensing Reconstruction of the Cosmic Microwave Background with Deep Neural Networks



Deblending galaxy superpositions with branched generative adversarial networks

David M. Reiman^{1*}[†], Brett E. Göhre¹^{‡§}

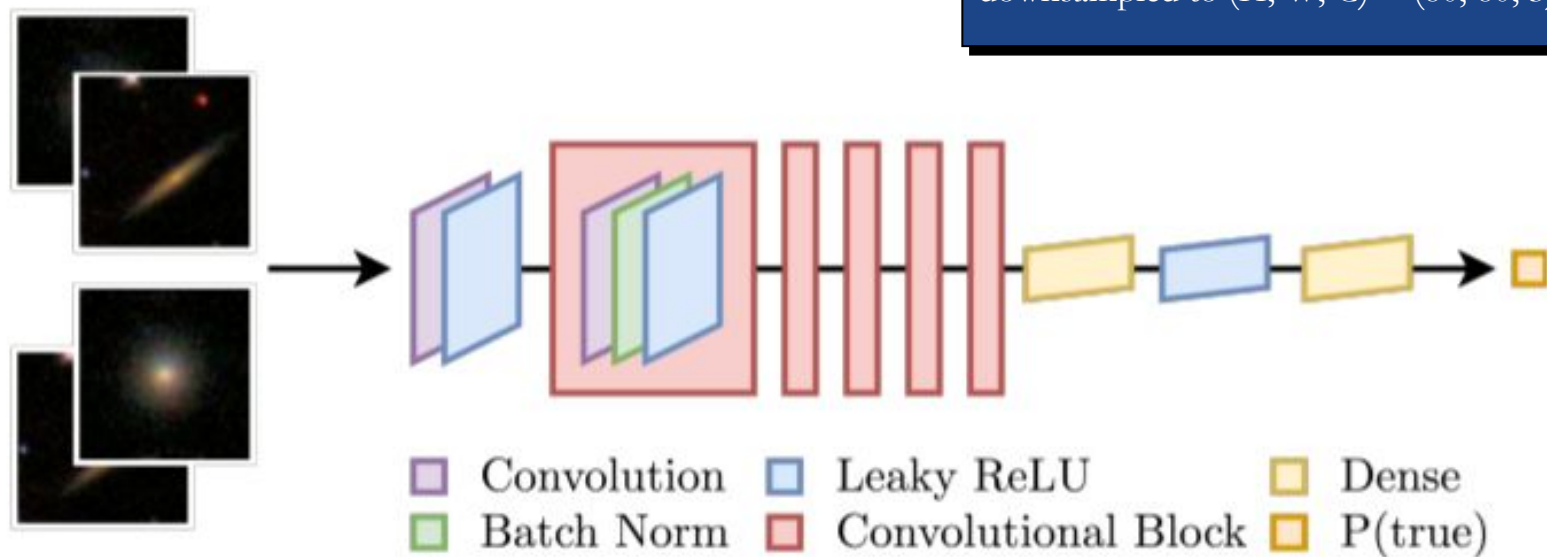


Deblending galaxy superpositions with branched generative adversarial networks

David M. Reiman¹^{★†}, Brett E. Göhre¹^{‡§}

¹Department of Physics, University of California Santa Cruz, 1156 High Street, Santa Cruz, California 95064,

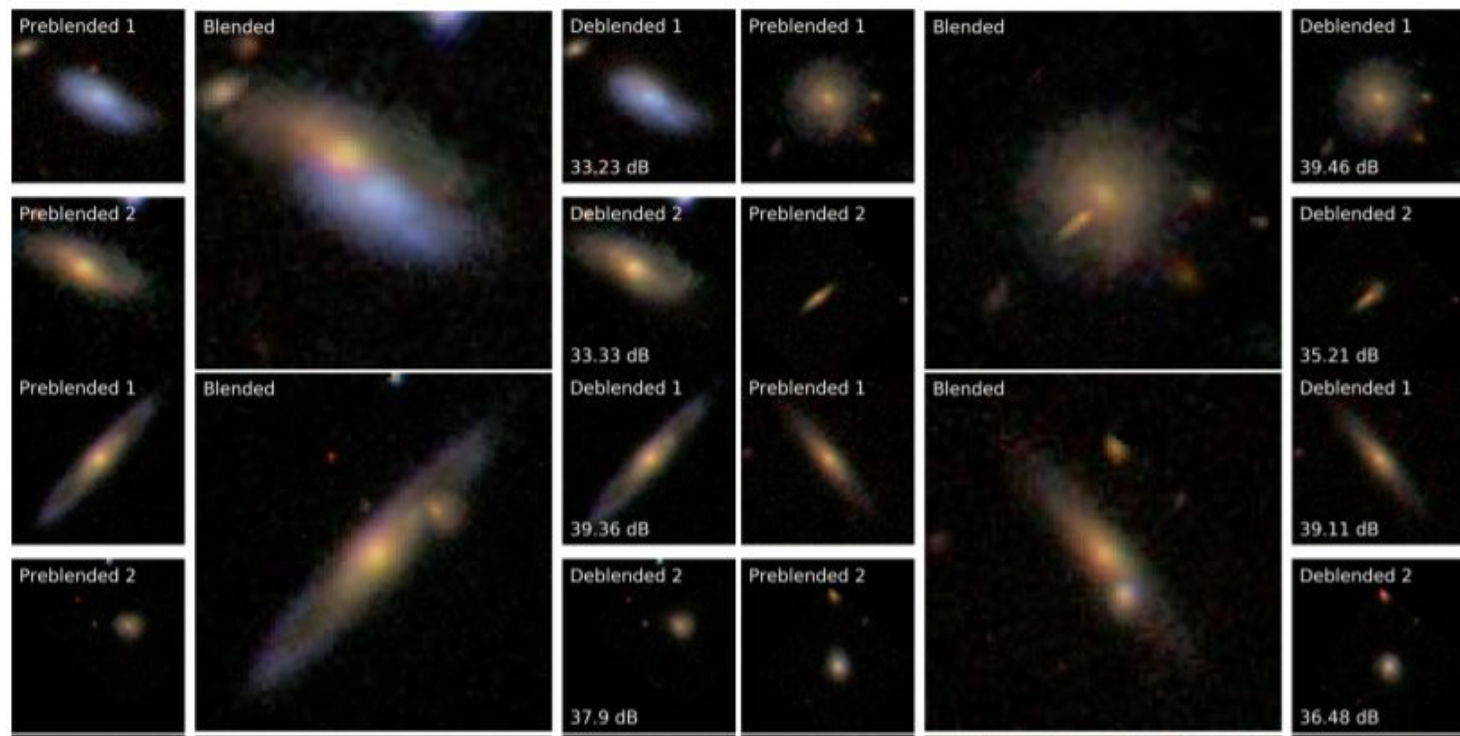
The authors used 141,553 images available from Galaxy Zoo. They used 3 bands in a JPEG format and therefore has a bit depth of 8 bits per color channel with dimension $(H, W, C) = (424, 424, 3)$ where H , W and C are the height, width and channel dimensions, respectively and were cropped and downsampled to $(H, W, C) = (80, 80, 3)$.



Deblending galaxy superpositions with branched generative adversarial networks

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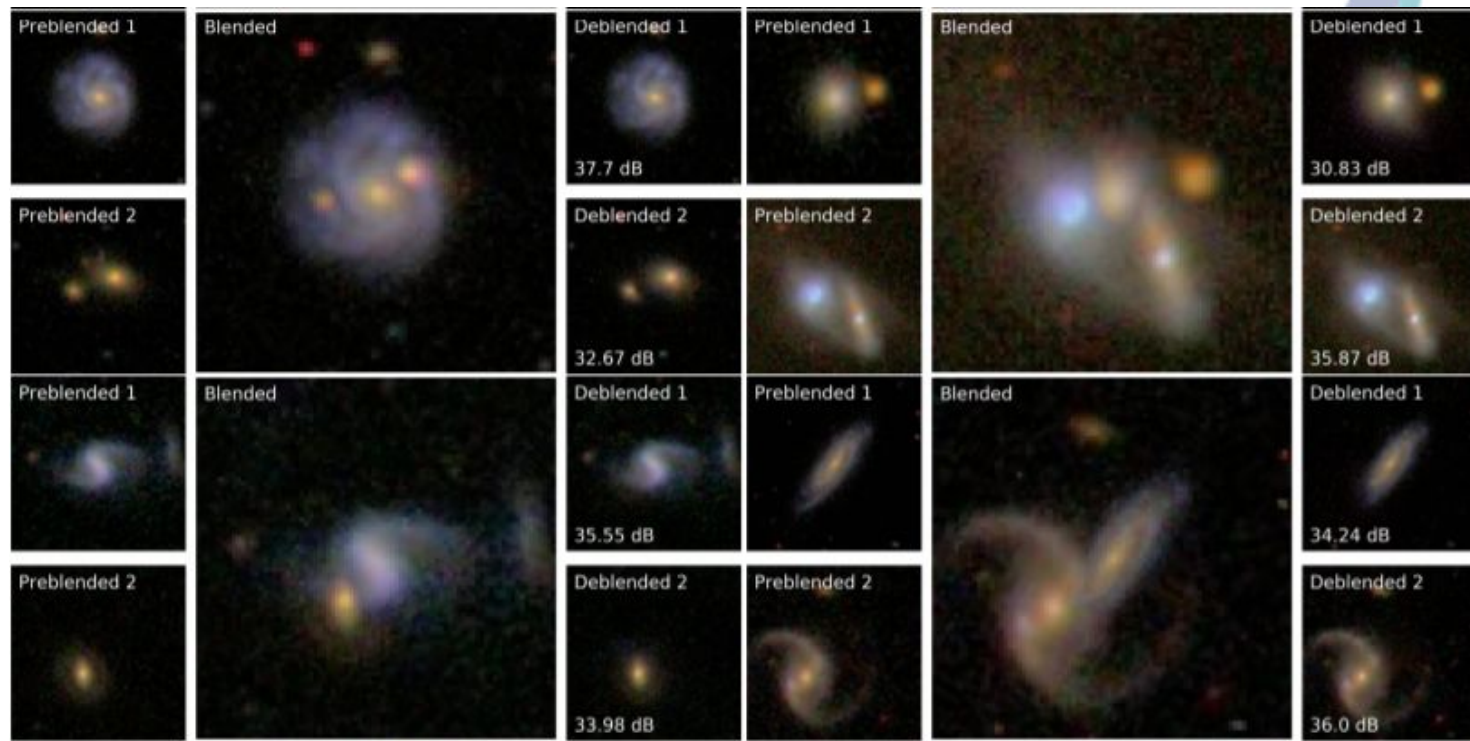
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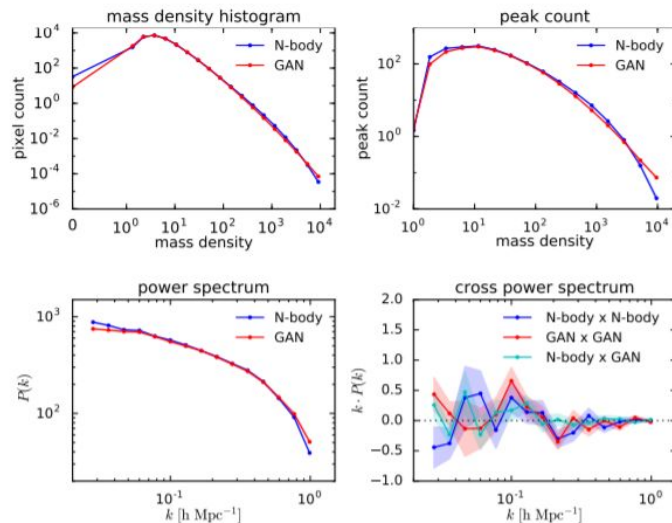
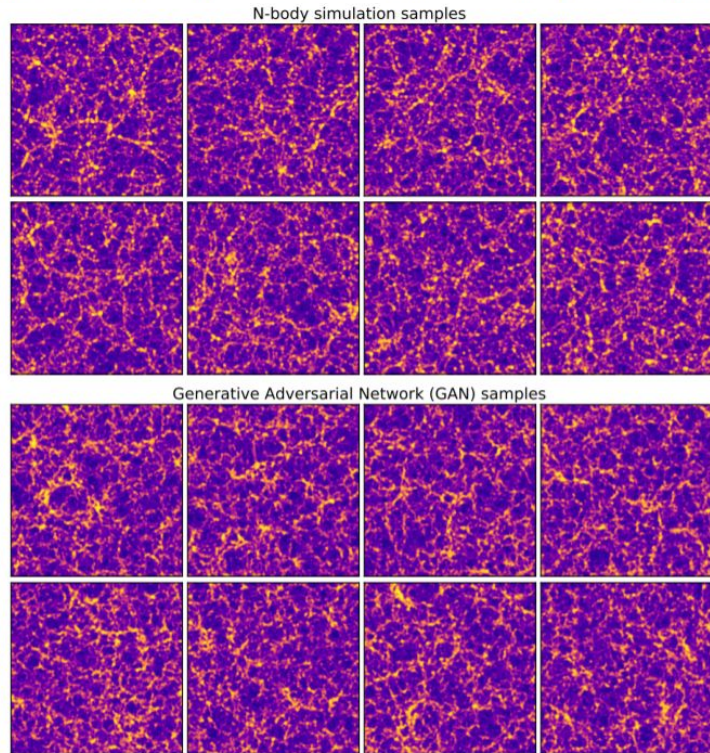


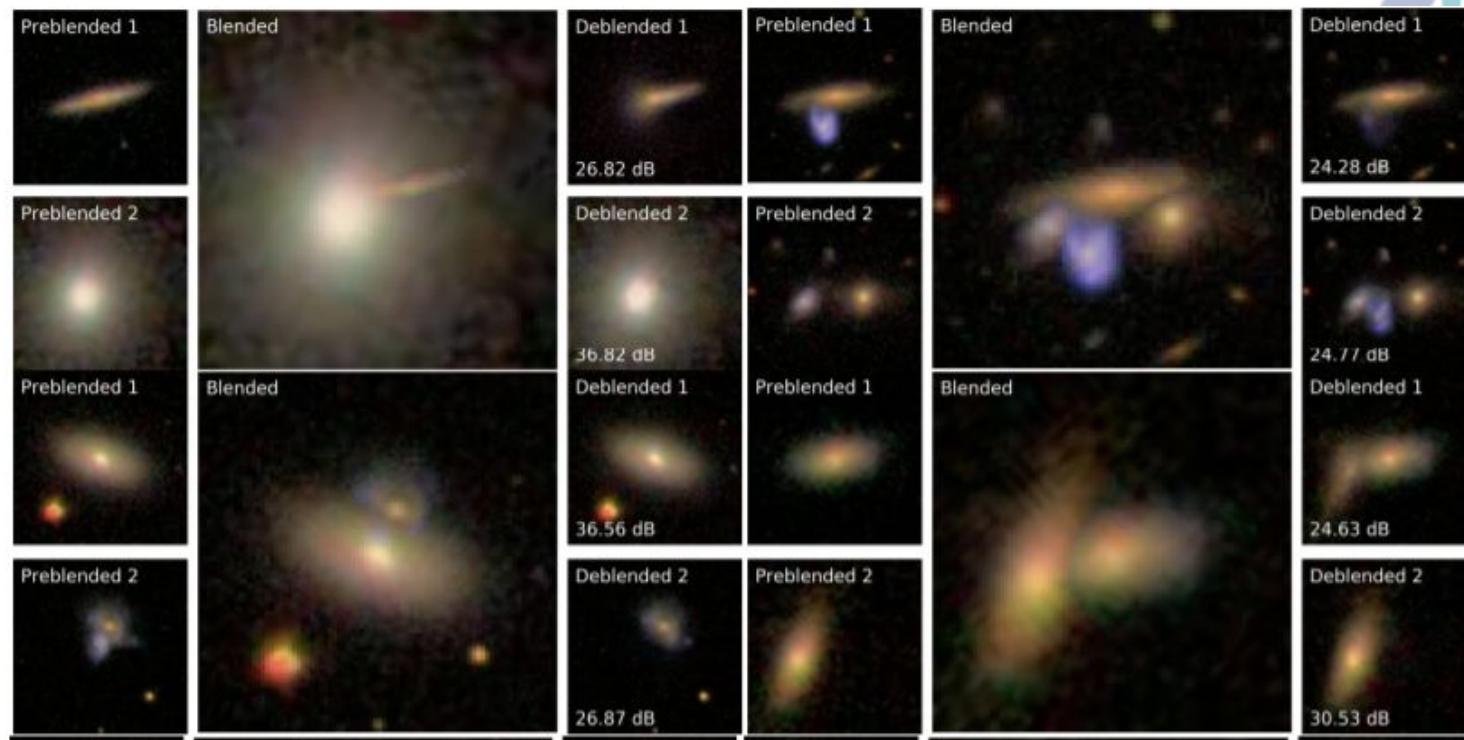
Figure 2: Comparison of summary statistics between N-body and GAN simulations, for box size of 500 Mpc. The statistics are: mass density histogram (upper left), peak count (upper right), power spectrum of 2D images (lower left) and cross power spectrum (lower right). The cross power spectrum is calculated between pairs N-body images (blue points), between pairs of GAN images (red points), and between pairs consisting of one GAN and one N-body image (cyan points). The power spectra are shown in units of h^{-1} Mpc, where $h = H_0/100$ corresponds to the Hubble parameter. The standard errors on the mean of the shown with a shaded region, and are too small to be seen for the first three panels.

Deblending galaxy superpositions with branched generative adversarial networks

David M. Reiman¹^{✶†}, Brett E. Göhre¹^{‡§}

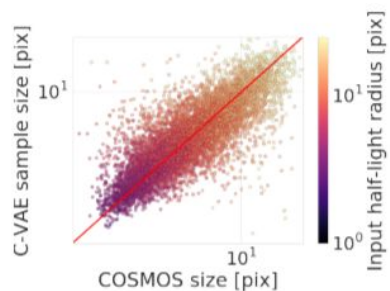
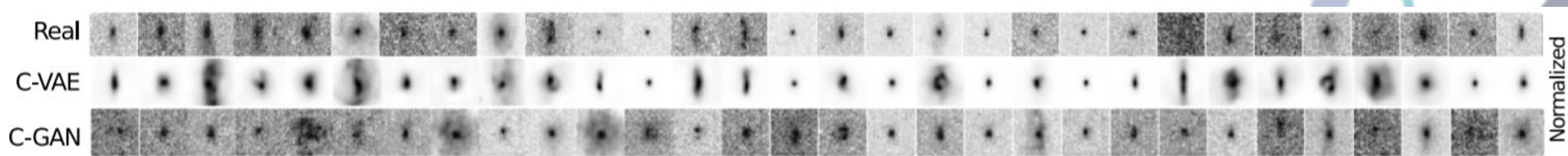
¹Department of Physics, University of California Santa Cruz, 1156 High Street, Santa Cruz, California 95064, USA

Failures

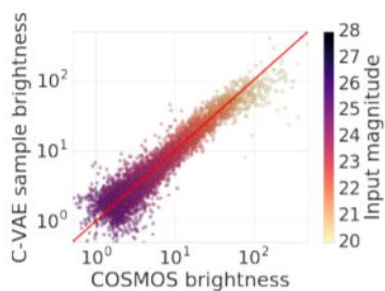


Enabling Dark Energy Science with Deep Generative Models of Galaxy Images

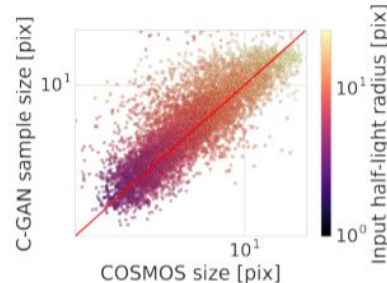
Siamak Ravanbakhsh¹, François Lanusse², Rachel Mandelbaum², Jeff Schneider¹, and Barnabás Póczos¹



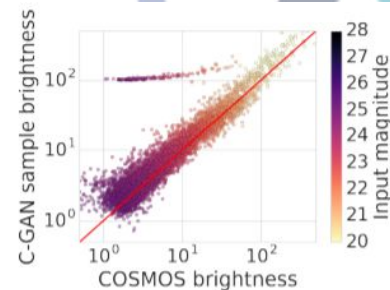
(a) Galaxy sizes



(b) Galaxy brightness

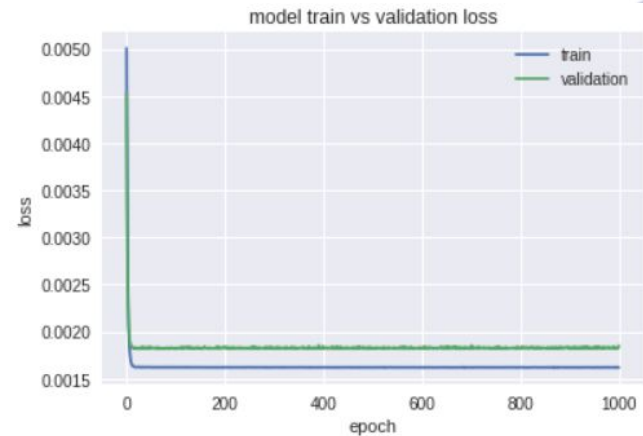
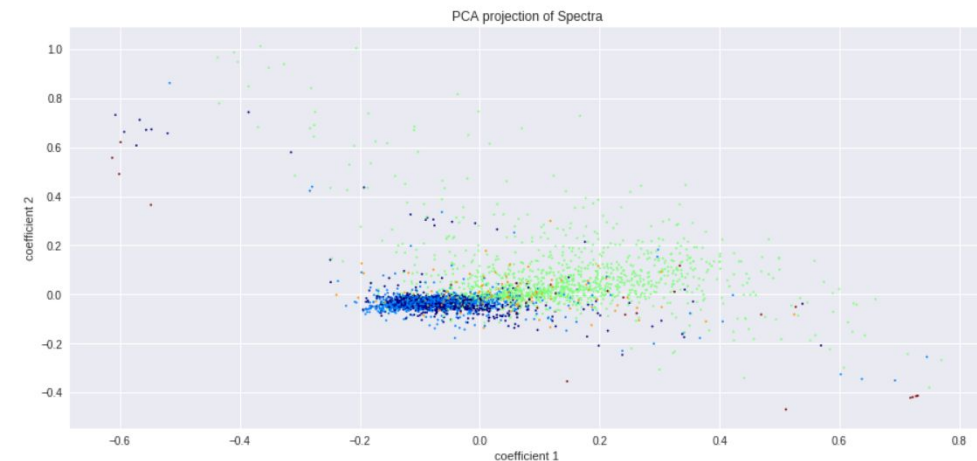


(a) Galaxy sizes

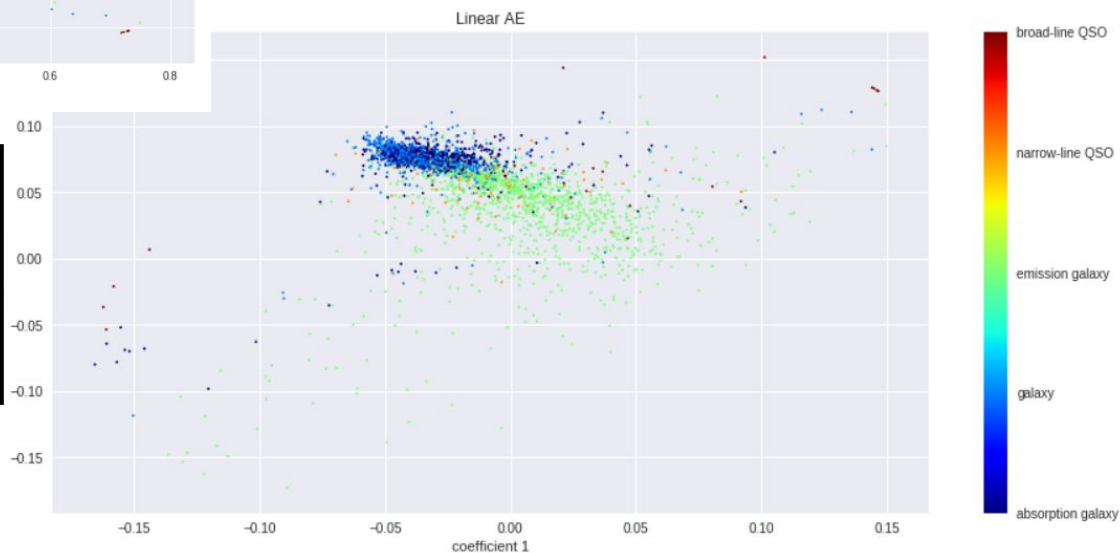


(b) Galaxy brightness

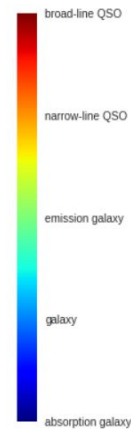
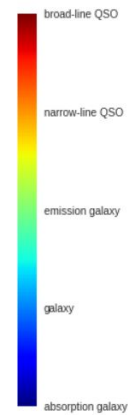
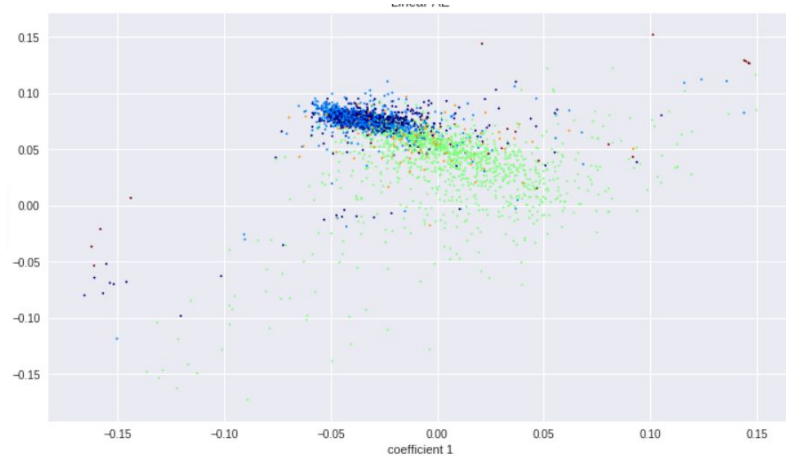
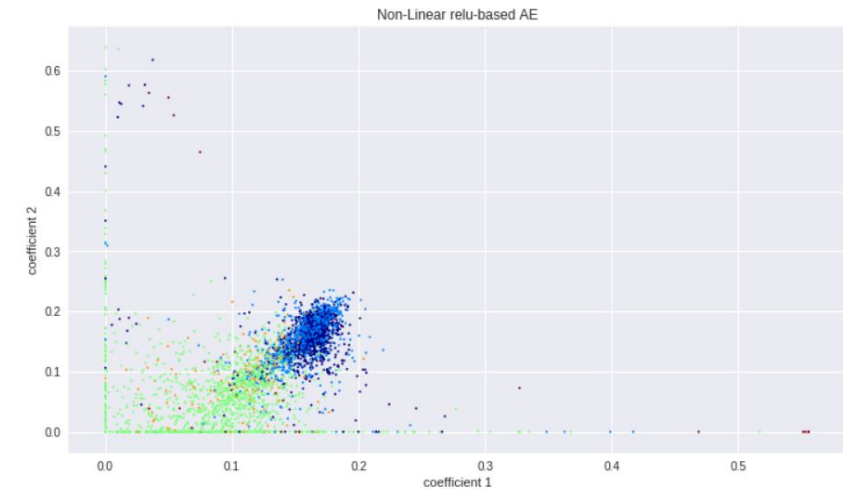
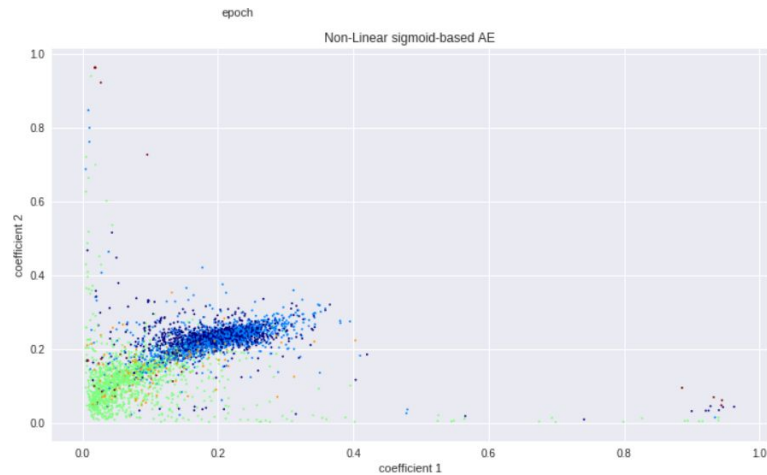
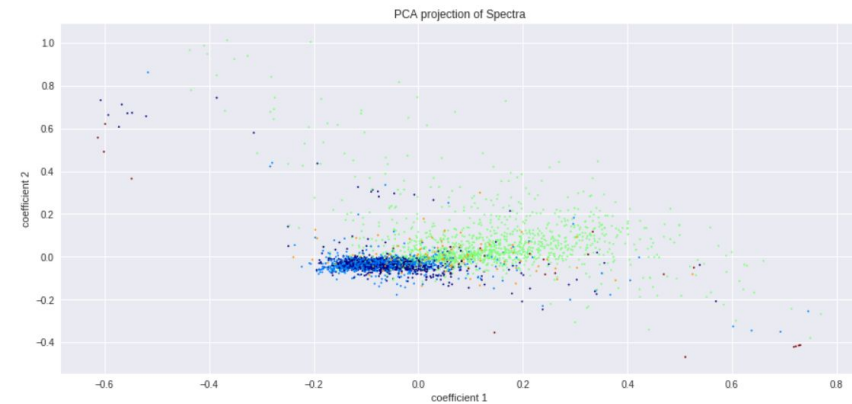
Example - PCA vs AE



Galaxy Spectra from SDSS available
from astroML example.
http://www.astroml.org/book_figures/chapter7/fig_PCA_LLE.html



Example - PCA vs AE





XII ESCOLA DO CBPF

22 de julho a 02 de agosto de 2019

Inteligência Artificial Utilizando *Deep Learning* e Aplicações em Física



**Márcio Portes
de Albuquerque
(CBPF)**



**Clécio R. De Bom
(CBPF/CEFET-RJ)**



**Elisangela L. Faria
(CBPF)**

Uma introdução conceitual de aprendizado de máquina e uma abordagem prática em aplicações de redes neurais profundas com foco em aplicações científicas e tecnológicas.

